
mhd-hermes Documentation

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Contents:

INSTALLATION INSTRUCTIONS

Install `hermes2d`, so that you can `import hermes2d` from Python:

```
In [1]: import hermes2d
```

```
In [2]:
```

Once this works, then just run:

```
cmake .  
make
```

and that's it (cmake will ask the `hermes2d` module where all the `*.h` and `*.pxd` files are).

MHD EQUATIONS

2.1 Introduction

The magnetohydrodynamics (MHD) equations are:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0 \quad (2.1)$$

$$\rho \left(\frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v} \cdot \nabla) \mathbf{v} \right) = -\nabla p + \frac{1}{\mu} (\nabla \times \mathbf{B}) \times \mathbf{B} + \rho \mathbf{g} \quad (2.2)$$

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{v} \times \mathbf{B}) + \eta \nabla^2 \mathbf{B} \quad (2.3)$$

$$\nabla \cdot \mathbf{B} = 0 \quad (2.4)$$

assuming η is constant. See the next section for a derivation. We can now apply the following identities (we use the fact that $\nabla \cdot \mathbf{B} = 0$):

$$\begin{aligned} [(\nabla \times \mathbf{B}) \times \mathbf{B}]_i &= \varepsilon_{ijk} (\nabla \times \mathbf{B})_j B_k = \varepsilon_{ijk} \varepsilon_{jlm} (\partial_l B_m) B_k = (\delta_{kl} \delta_{im} - \delta_{km} \delta_{il}) (\partial_l B_m) B_k = \\ &= (\partial_k B_i) B_k - (\partial_i B_k) B_k = \left[(\mathbf{B} \cdot \nabla) \mathbf{B} - \frac{1}{2} \nabla |\mathbf{B}|^2 \right]_i \end{aligned}$$

$$(\nabla \times \mathbf{B}) \times \mathbf{B} = (\mathbf{B} \cdot \nabla) \mathbf{B} - \frac{1}{2} \nabla |\mathbf{B}|^2 = (\mathbf{B} \cdot \nabla) \mathbf{B} + \mathbf{B} (\nabla \cdot \mathbf{B}) - \frac{1}{2} \nabla |\mathbf{B}|^2 = \nabla \cdot (\mathbf{B} \mathbf{B}^T) - \frac{1}{2} \nabla |\mathbf{B}|^2$$

$$\nabla \times (\mathbf{v} \times \mathbf{B}) = (\mathbf{B} \cdot \nabla) \mathbf{v} - \mathbf{B} (\nabla \cdot \mathbf{v}) + \mathbf{v} (\nabla \cdot \mathbf{B}) - (\mathbf{v} \cdot \nabla) \mathbf{B} = \nabla \cdot (\mathbf{B} \mathbf{v}^T - \mathbf{v} \mathbf{B}^T)$$

$$\nabla \cdot (\rho \mathbf{v} \mathbf{v}^T) = (\nabla \cdot (\rho \mathbf{v})) \mathbf{v} + \rho (\mathbf{v} \cdot \nabla) \mathbf{v} = -\mathbf{v} \frac{\partial \rho}{\partial t} + \rho (\mathbf{v} \cdot \nabla) \mathbf{v}$$

So the MHD equations can alternatively be written as:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0 \quad (2.5)$$

$$\frac{\partial \rho \mathbf{v}}{\partial t} + \nabla \cdot (\rho \mathbf{v} \mathbf{v}^T) = -\nabla p + \frac{1}{\mu} \left(\nabla \cdot (\mathbf{B} \mathbf{B}^T) - \frac{1}{2} \nabla |\mathbf{B}|^2 \right) + \rho \mathbf{g} \quad (2.6)$$

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \cdot (\mathbf{B} \mathbf{v}^T - \mathbf{v} \mathbf{B}^T) + \eta \nabla^2 \mathbf{B} \quad (2.7)$$

$$\nabla \cdot \mathbf{B} = 0 \quad (2.8)$$

One can also introduce a new variable $p^* = p + \frac{1}{2} \nabla |\mathbf{B}|^2$, that simplifies (2.6) a bit.

2.2 Derivation

The above equations can easily be derived. We have the continuity equation:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0$$

Navier-Stokes equations (momentum equation) with the Lorentz force on the right-hand side:

$$\rho \left(\frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v} \cdot \nabla) \mathbf{v} \right) = -\nabla p + \mathbf{j} \times \mathbf{B} + \rho \mathbf{g}$$

where the current density \mathbf{j} is given by the Maxwell equation (we neglect the displacement current $\frac{\partial \mathbf{E}}{\partial t}$):

$$\mathbf{j} = \frac{1}{\mu} \nabla \times \mathbf{B}$$

and the Lorentz force:

$$\frac{1}{\sigma} \mathbf{j} = \mathbf{E} + \mathbf{v} \times \mathbf{B}$$

from which we eliminate \mathbf{E} :

$$\mathbf{E} = -\mathbf{v} \times \mathbf{B} + \frac{1}{\sigma} \mathbf{j} = -\mathbf{v} \times \mathbf{B} + \frac{1}{\sigma \mu} \nabla \times \mathbf{B}$$

and put it into the Maxwell equation:

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E}$$

so we get:

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{v} \times \mathbf{B}) - \nabla \times \left(\frac{1}{\sigma \mu} \nabla \times \mathbf{B} \right)$$

assuming the magnetic diffusivity $\eta = \frac{1}{\sigma \mu}$ is constant, we get:

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{v} \times \mathbf{B}) - \eta \nabla \times (\nabla \times \mathbf{B}) = \nabla \times (\mathbf{v} \times \mathbf{B}) + \eta (\nabla^2 \mathbf{B} - \nabla(\nabla \cdot \mathbf{B})) = \nabla \times (\mathbf{v} \times \mathbf{B}) + \eta \nabla^2 \mathbf{B}$$

where we used the Maxwell equation:

$$\nabla \cdot \mathbf{B} = 0$$

2.3 Finite Element Formulation

We solve the following ideal MHD equations (we use $p^* = p + \frac{1}{2} \nabla |\mathbf{B}|^2$, but we drop the star):

$$\frac{\partial \mathbf{u}}{\partial t} + (\mathbf{u} \cdot \nabla) \mathbf{u} - (\mathbf{B} \cdot \nabla) \mathbf{B} + \nabla p = 0 \tag{2.9}$$

$$\frac{\partial \mathbf{B}}{\partial t} + (\mathbf{u} \cdot \nabla) \mathbf{B} - (\mathbf{B} \cdot \nabla) \mathbf{u} = 0 \tag{2.10}$$

$$\nabla \cdot \mathbf{u} = 0 \tag{2.11}$$

$$\nabla \cdot \mathbf{B} = 0 \quad (2.12)$$

If the equation (2.12) is satisfied initially, then it is satisfied all the time, as can be easily proved by applying a divergence to the Maxwell equation $\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E}$ and we get $\frac{\partial}{\partial t}(\nabla \cdot \mathbf{B}) = 0$, so $\nabla \cdot \mathbf{B}$ is constant, independent of time. As a consequence, we are essentially only solving equations (2.9), (2.10) and (2.11), which consist of 5 equations for 5 unknowns (components of \mathbf{u} , p and \mathbf{B}).

We discretize in time by introducing a small time step τ and we also linearize the convective terms:

$$\frac{\mathbf{u}^n - \mathbf{u}^{n-1}}{\tau} + (\mathbf{u}^{n-1} \cdot \nabla) \mathbf{u}^n - (\mathbf{B}^{n-1} \cdot \nabla) \mathbf{B}^n + \nabla p = 0 \quad (2.13)$$

$$\frac{\mathbf{B}^n - \mathbf{B}^{n-1}}{\tau} + (\mathbf{u}^{n-1} \cdot \nabla) \mathbf{B}^n - (\mathbf{B}^{n-1} \cdot \nabla) \mathbf{u}^n = 0 \quad (2.14)$$

$$\nabla \cdot \mathbf{u}^n = 0 \quad (2.15)$$

Testing (2.13) by the test functions (v_1, v_2) , (2.14) by the functions (C_1, C_2) and (2.15) by the test function q , we obtain the following weak formulation:

$$\begin{aligned} \int_{\Omega} \frac{u_1 v_1}{\tau} + (\mathbf{u}^{n-1} \cdot \nabla) u_1 v_1 - (\mathbf{B}^{n-1} \cdot \nabla) B_1 v_1 - p \frac{\partial v_1}{\partial x} \, d\mathbf{x} &= \int_{\Omega} \frac{u_1^{n-1} v_1}{\tau} \, d\mathbf{x} \\ \int_{\Omega} \frac{u_2 v_2}{\tau} + (\mathbf{u}^{n-1} \cdot \nabla) u_2 v_2 - (\mathbf{B}^{n-1} \cdot \nabla) B_2 v_2 - p \frac{\partial v_2}{\partial y} \, d\mathbf{x} &= \int_{\Omega} \frac{u_2^{n-1} v_2}{\tau} \, d\mathbf{x} \end{aligned} \quad (2.16)$$

$$\begin{aligned} \int_{\Omega} \frac{B_1 C_1}{\tau} + (\mathbf{u}^{n-1} \cdot \nabla) B_1 C_1 - (\mathbf{B}^{n-1} \cdot \nabla) u_1 C_1 \, d\mathbf{x} &= \int_{\Omega} \frac{B_1^{n-1} C_1}{\tau} \, d\mathbf{x} \\ \int_{\Omega} \frac{B_2 C_2}{\tau} + (\mathbf{u}^{n-1} \cdot \nabla) B_2 C_2 - (\mathbf{B}^{n-1} \cdot \nabla) u_2 C_2 \, d\mathbf{x} &= \int_{\Omega} \frac{B_2^{n-1} C_2}{\tau} \, d\mathbf{x} \end{aligned} \quad (2.17)$$

$$\int_{\Omega} \frac{\partial u_1}{\partial x} q + \frac{\partial u_2}{\partial y} q \, d\mathbf{x} = 0 \quad (2.18)$$

To better understand the structure of these equations, we write it using bilinear and linear forms, as well as take into account the symmetries of the forms. Then we get a particularly simple structure:

$$\begin{array}{rclcl} +A(u_1, v_1) & & -X(p, v_1) & -B(B_1, v_1) & = l_1(v_1) \\ & +A(u_2, v_2) & -Y(p, v_2) & & = l_2(v_2) \\ +X(q, u_1) & +Y(q, u_2) & & & = 0 \\ -B(u_1, C_1) & & & +A(B_1, C_1) & = l_4(C_1) \\ & -B(u_2, C_2) & & +A(B_2, C_2) & = l_5(C_2) \end{array}$$

where:

$$\begin{aligned}
 A(u, v) &= \int_{\Omega} \frac{uv}{\tau} + (\mathbf{u}^{n-1} \cdot \nabla)uv \, d\mathbf{x} \\
 B(u, v) &= \int_{\Omega} (\mathbf{B}^{n-1} \cdot \nabla)uv \, d\mathbf{x} \\
 X(u, v) &= \int_{\Omega} u \frac{\partial v}{\partial x} \, d\mathbf{x} \\
 Y(u, v) &= \int_{\Omega} u \frac{\partial v}{\partial y} \, d\mathbf{x} \\
 l_1(v) &= \int_{\Omega} \frac{u_1^{n-1}v}{\tau} \, d\mathbf{x} \\
 l_2(v) &= \int_{\Omega} \frac{u_2^{n-1}v}{\tau} \, d\mathbf{x} \\
 l_4(v) &= \int_{\Omega} \frac{B_1^{n-1}v}{\tau} \, d\mathbf{x} \\
 l_5(v) &= \int_{\Omega} \frac{B_2^{n-1}v}{\tau} \, d\mathbf{x}
 \end{aligned}$$

E.g. there are only 4 distinct bilinear forms. Schematically we can visualize the structure by:

A		-X	-B	
X	A	-Y		-B
-B	-B		A	A

In order to solve it with Hermes, we first need to write it in the block form:

$$\begin{aligned}
 a_{11}(u_1, v_1) + a_{12}(u_2, v_1) + a_{13}(p, v_1) + a_{14}(B_1, v_1) + a_{15}(B_2, v_1) &= l_1(v_1) \\
 a_{21}(u_1, v_2) + a_{22}(u_2, v_2) + a_{23}(p, v_2) + a_{24}(B_1, v_2) + a_{25}(B_2, v_2) &= l_2(v_2) \\
 a_{31}(u_1, q) + a_{32}(u_2, q) + a_{33}(p, q) + a_{34}(B_1, q) + a_{35}(B_2, q) &= l_3(q) \\
 a_{41}(u_1, C_1) + a_{42}(u_2, C_1) + a_{43}(p, C_1) + a_{44}(B_1, C_1) + a_{45}(B_2, C_1) &= l_4(C_1) \\
 a_{51}(u_1, C_2) + a_{52}(u_2, C_2) + a_{53}(p, C_2) + a_{54}(B_1, C_2) + a_{55}(B_2, C_2) &= l_5(C_2)
 \end{aligned}$$

comparing to the above, we get the following nonzero forms:

$$\begin{aligned}
 a_{11}(u_1, v_1) + 0 + a_{13}(p, v_1) + a_{14}(B_1, v_1) + 0 &= l_1(v_1) \\
 0 + a_{22}(u_2, v_2) + a_{23}(p, v_2) + 0 + a_{25}(B_2, v_2) &= l_2(v_2) \\
 a_{31}(u_1, q) + a_{32}(u_2, q) + 0 + 0 + 0 &= 0 \\
 a_{41}(u_1, C_1) + 0 + 0 + a_{44}(B_1, C_1) + 0 &= l_4(C_1) \\
 0 + a_{52}(u_2, C_2) + 0 + 0 + a_{55}(B_2, C_2) &= l_5(C_2)
 \end{aligned}$$

where:

$$\begin{aligned}a_{11}(u_1, v_1) &= A(u_1, v_1) \\a_{22}(u_2, v_2) &= A(u_2, v_2) \\a_{44}(B_1, C_1) &= A(B_1, C_1) \\a_{55}(B_2, C_2) &= A(B_2, C_2) \\a_{13}(p, v_1) &= -X(p, v_1) \\a_{31}(u_1, q) &= X(q, u_1) \\a_{23}(p, v_2) &= -Y(p, v_2) \\a_{32}(u_2, q) &= Y(q, u_2) \\a_{14}(B_1, v_1) &= -B(B_1, v_1) \\a_{41}(u_1, C_1) &= -B(u_1, C_1) \\a_{25}(B_2, v_2) &= -B(B_2, v_2) \\a_{52}(u_2, C_2) &= -B(u_2, C_2)\end{aligned}$$

and l_1, \dots, l_5 are the same as above.

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